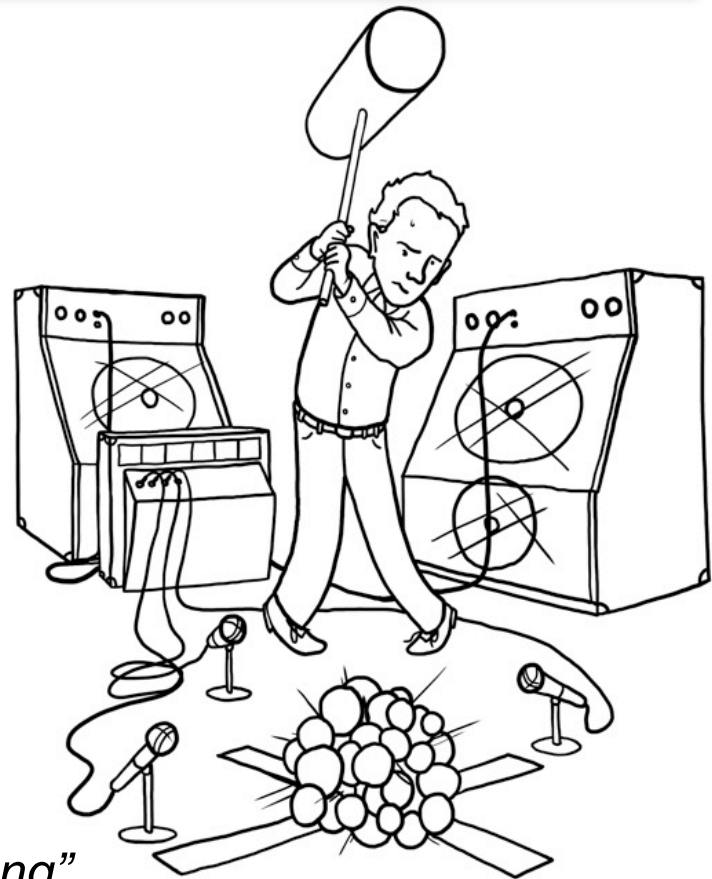


# Status and Prospects in Heavy Flavor Probes

Ágnes Mócsy  
Pratt Institute, Brooklyn, NY

**Pratt**



*“The Sound of the Little Bang”*

*Illustration by Alex Doig (Pratt Institute)*

# Heavy flavor / Quarkonium Wiki

Quarkonium in media - QPG

https://wiki.bnl.gov/qpg/index.php/Quarkonium\_in\_media#Cold\_Nuclear\_Matter\_28CNM.29\_Effect

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## Quarkonium in media

Contents [hide]

- 1 Introduction
- 2 Quarkonium properties and quarkonium spectral functions in finite temperature QCD
  - 2.1 Color screening in high temperature QCD
  - 2.2 Quarkonia properties at finite temperature : potential models, effective field theory, sum rules
  - 2.3 Spectral functions and lattice QCD
- 3 Cold Nuclear Matter (CNM) Effects on quarkonium production
- 4 Quarkonium production in AA collisions
  - 4.1 Mechanism of quarkonium production
  - 4.2 Experimental results
  - 4.3 Dynamical models for quarkonium productions in HI collisions
- 5 Useful Links
  - 5.1 Forums
  - 5.2 QWG document
  - 5.3 Recent review articles
  - 5.4 Upcoming meetings
  - 5.5 Past meetings

### Introduction

Quarkonia can provide a valuable tool for studying matter produced in heavy ion (HI) collisions, in particular they can probe deconfinement. In order to use quarkonia as probes the following prerequisites are needed

- 1) Understanding of quarkonium properties in thermal QGP
- 2) Understanding of cold nuclear matter effects
- 3) Dynamical models to relate information from thermal QCD to experimental data

These issues will be discussed below.

[edit]

<https://wiki.bnl.gov/qpg/index.php/Quarkonia>

# Heavy flavor / Quarkonium Wiki

Quarkonium in media - QPG

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## Quarkonium in media

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### Introduction

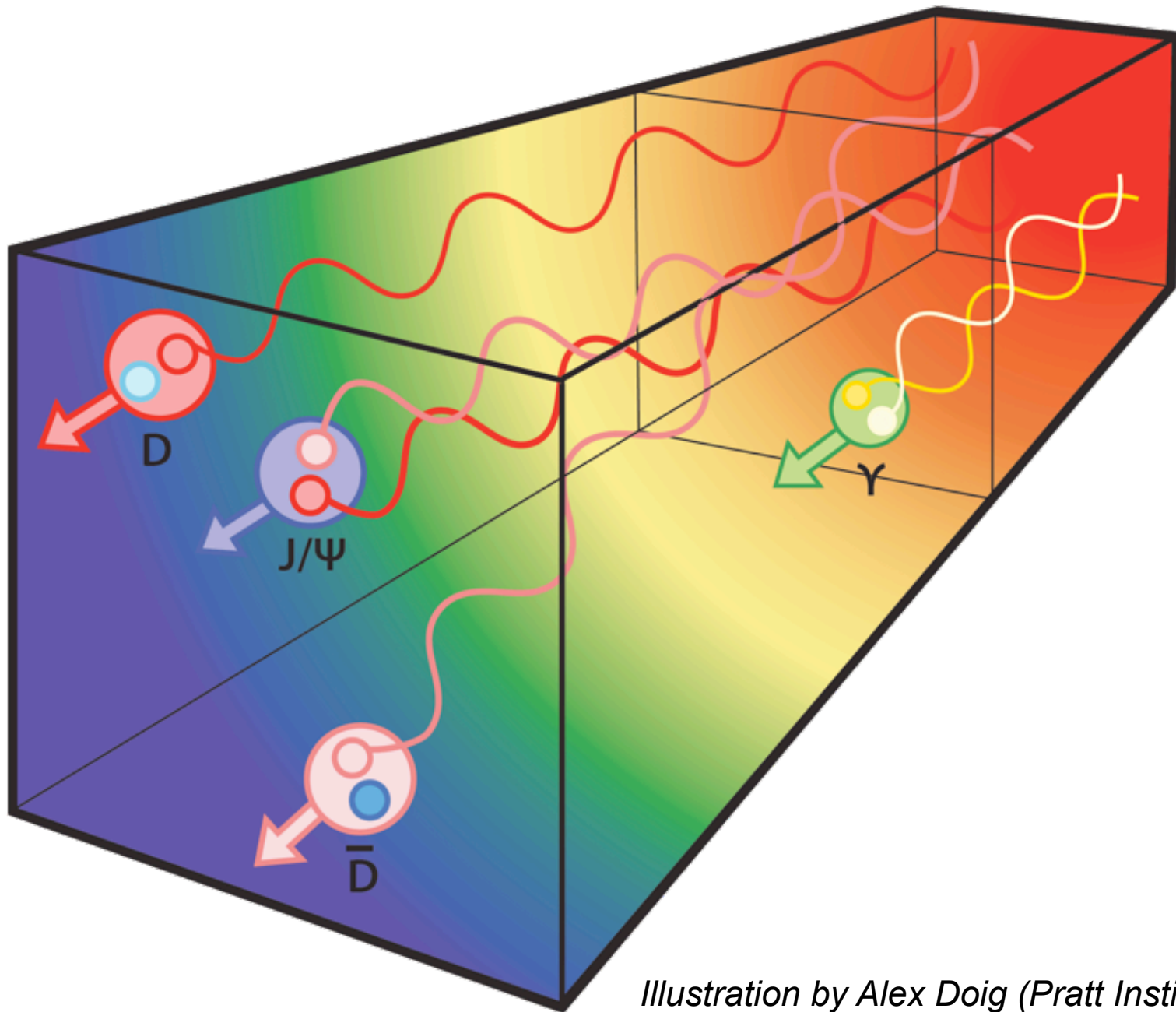
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[edit]

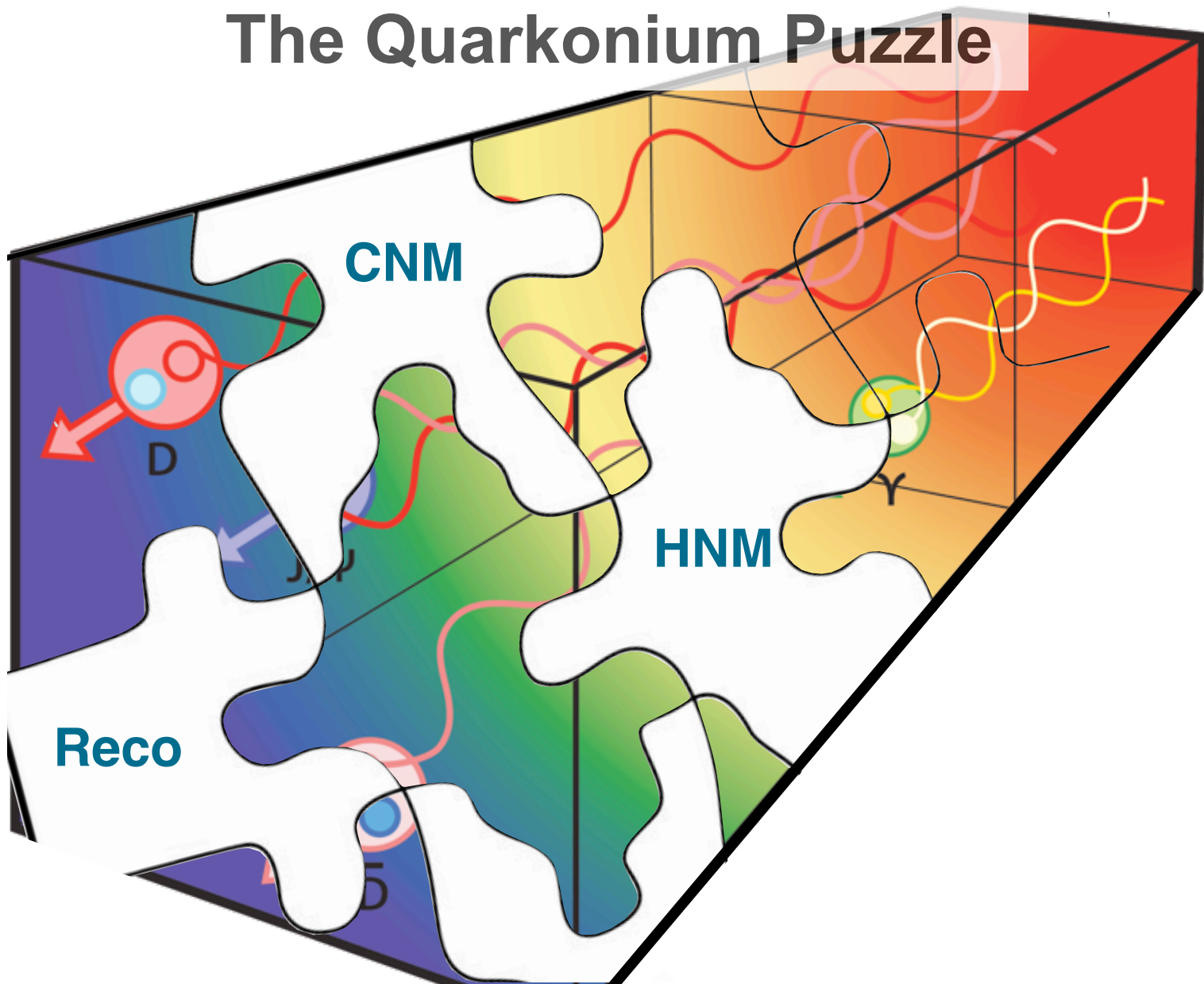
Contributions from everyone are welcome!

<https://wiki.bnl.gov/qpg/index.php/Quarkonia>



*Illustration by Alex Doig (Pratt Institute)*

# The Quarkonium Puzzle



*Illustration by Alex Doig (Pratt Institute)*

# Cold Nuclear Matter Effects

(Still is) A complicated problem

**Talks by Ramona Vogt, Tony Frawley,  
Alex Linden-Levy, Jamie Nagle**

# Where do we stand ?

“We are revisiting all the issues we thought were resolved 10 years ago”

Ramona Vogt  
(Jamie Nagle)



# Where do we stand ?

“We are revisiting all the issues we thought were resolved 10 years ago”

Ramona Vogt  
(Jamie Nagle)

Energy dependence of the absorption cross section  
*paper by Gavin and Vogt 2000*



## The problem:

The heavy ion  $J/\psi$  data alone have not taught us as much as we would like, because of serious uncertainties caused by:

- 1) Poorly known initial state effects at RHIC:
  - Break up cross section for collisions with nucleons.
  - Shadowing.
  - Other effects? Initial state energy loss?
- 2) Poorly known open charm production cross sections.

Thus the trade-off between coalescence and destruction is difficult to illuminate experimentally.

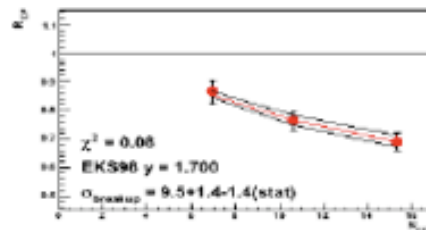
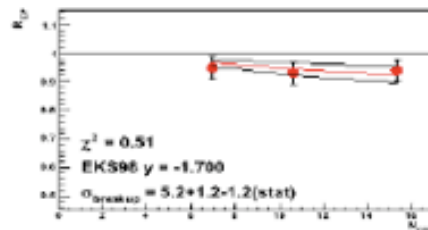
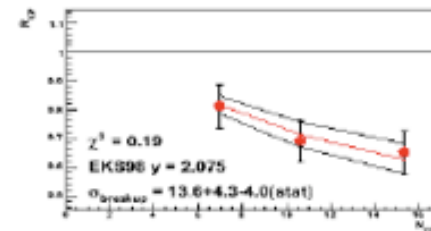
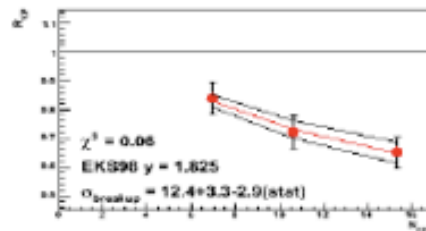
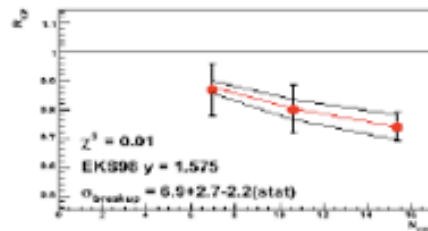
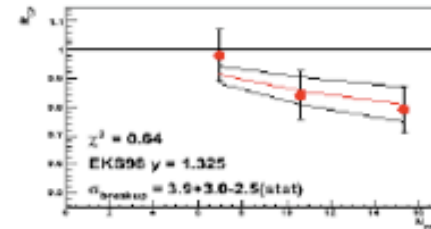
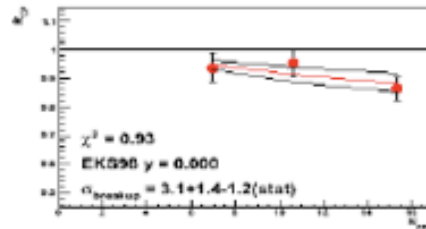
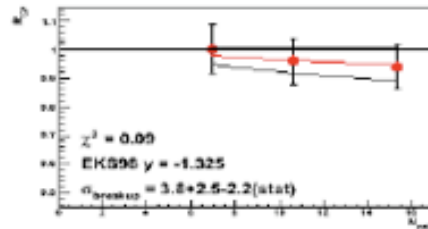
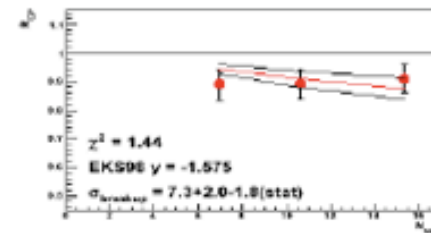
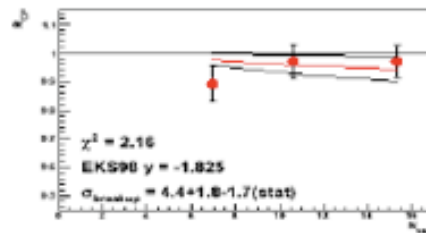
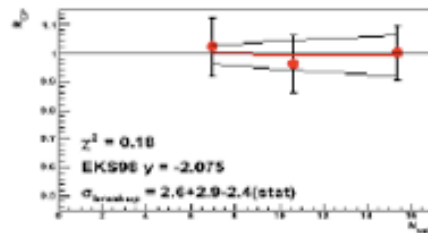
To try to make inroads on 1), we start from the most recent d+Au data set:

**– Run 8 d+Au**

and try to understand what the Au+Au  $R_{AA}$  would look like without hot nuclear matter effects.

**Talk by T Frawley**

# Fits to d+Au $R_{CP}$ – example for EKS98



Integrated for each  
muon arm

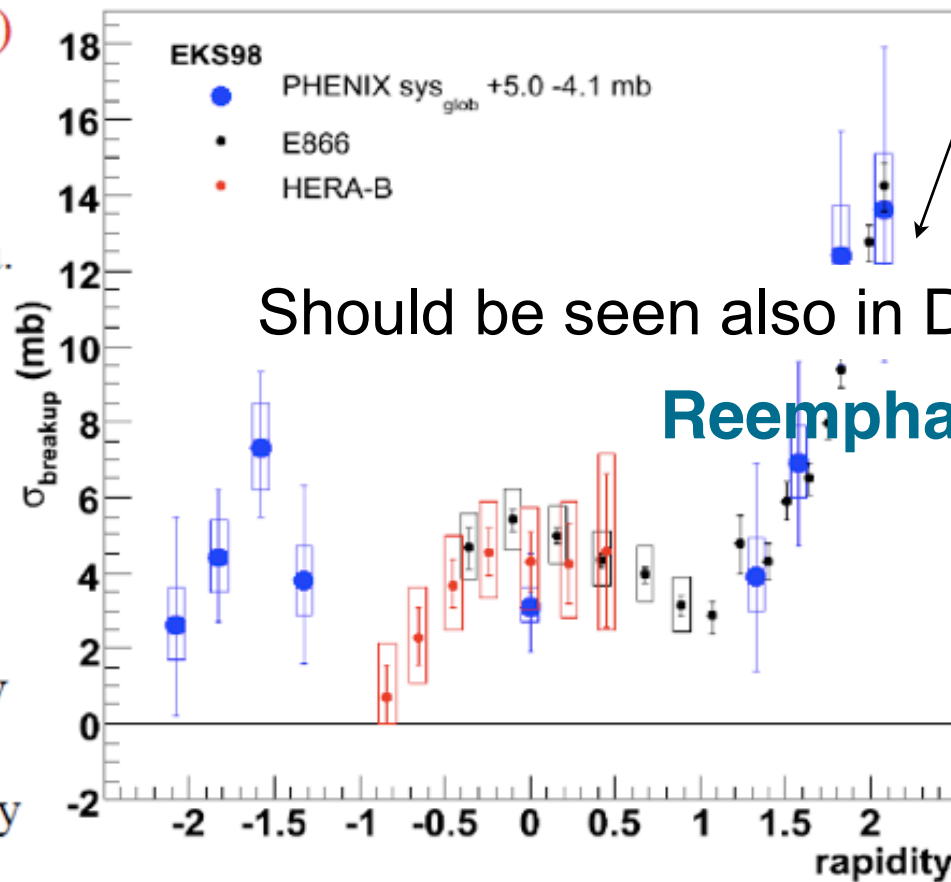
Talk by T Frawley

## Comparison with lower energy data – EKS98 fits

Lourenco, Vogt and Woehri (**JHEP 02 (2009) 014**) published the effective breakup cross section vs  $y$  from fits to E866 and HERA-B data.

Our results from 200 GeV are shown here compared with their results, both for the EKS98 case.

For  $y > 1.2$  the 200 GeV data follow the trend observed at lower energy remarkably closely!

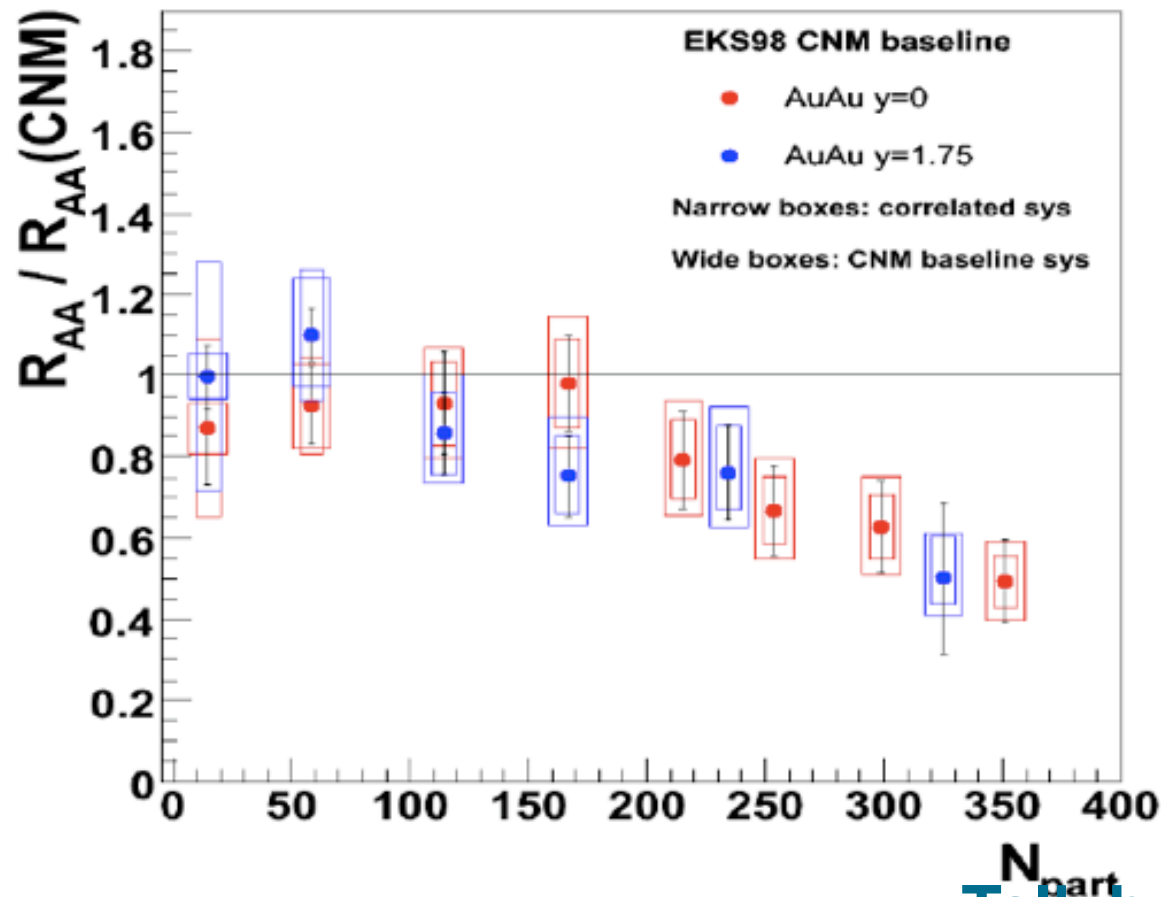


**Reemphasized by J. Nagle**  
check out Kopeliovics

**Talk by T Frawley**

## Results for Au+Au for the EKS98 case

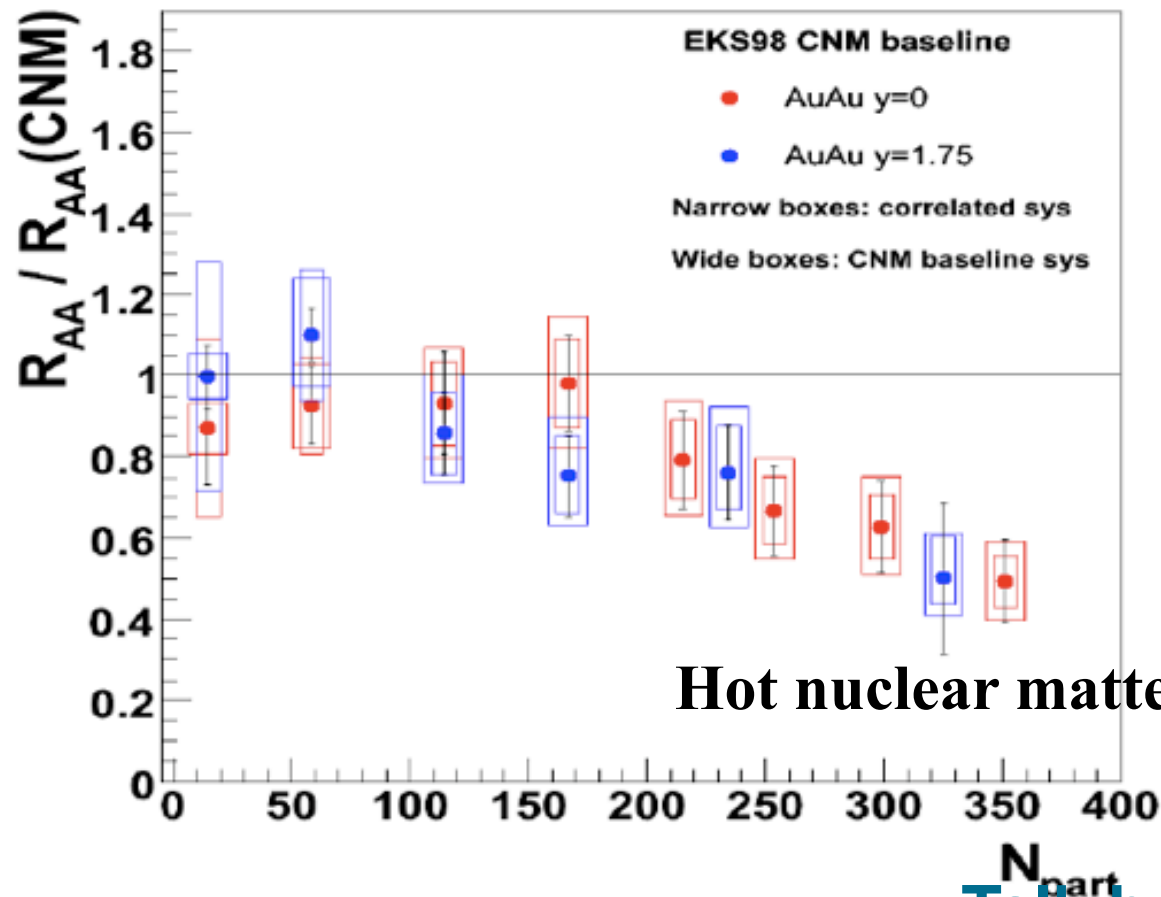
The suppression beyond CNM effects is found to be similar at  $y=0$  and at  $y=1.7$ . There is essentially **no** dependence on the shadowing model used to parameterize the d+Au  $R_{CP}$ .



Talk by T Frawley

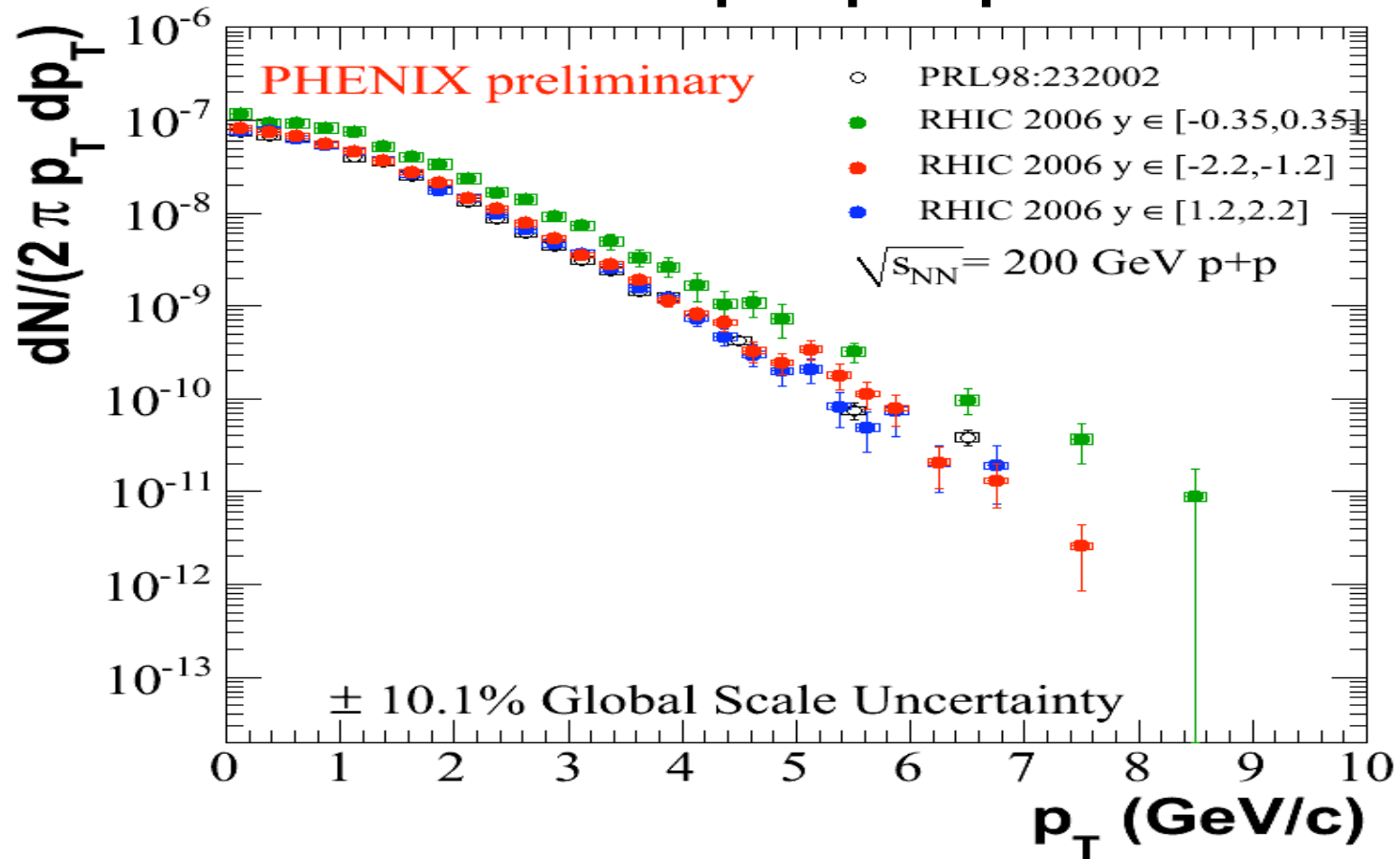
## Results for Au+Au for the EKS98 case

The suppression beyond CNM effects is found to be similar at  $y=0$  and at  $y=1.7$ . There is essentially **no** dependence on the shadowing model used to parameterize the d+Au  $R_{CP}$ .



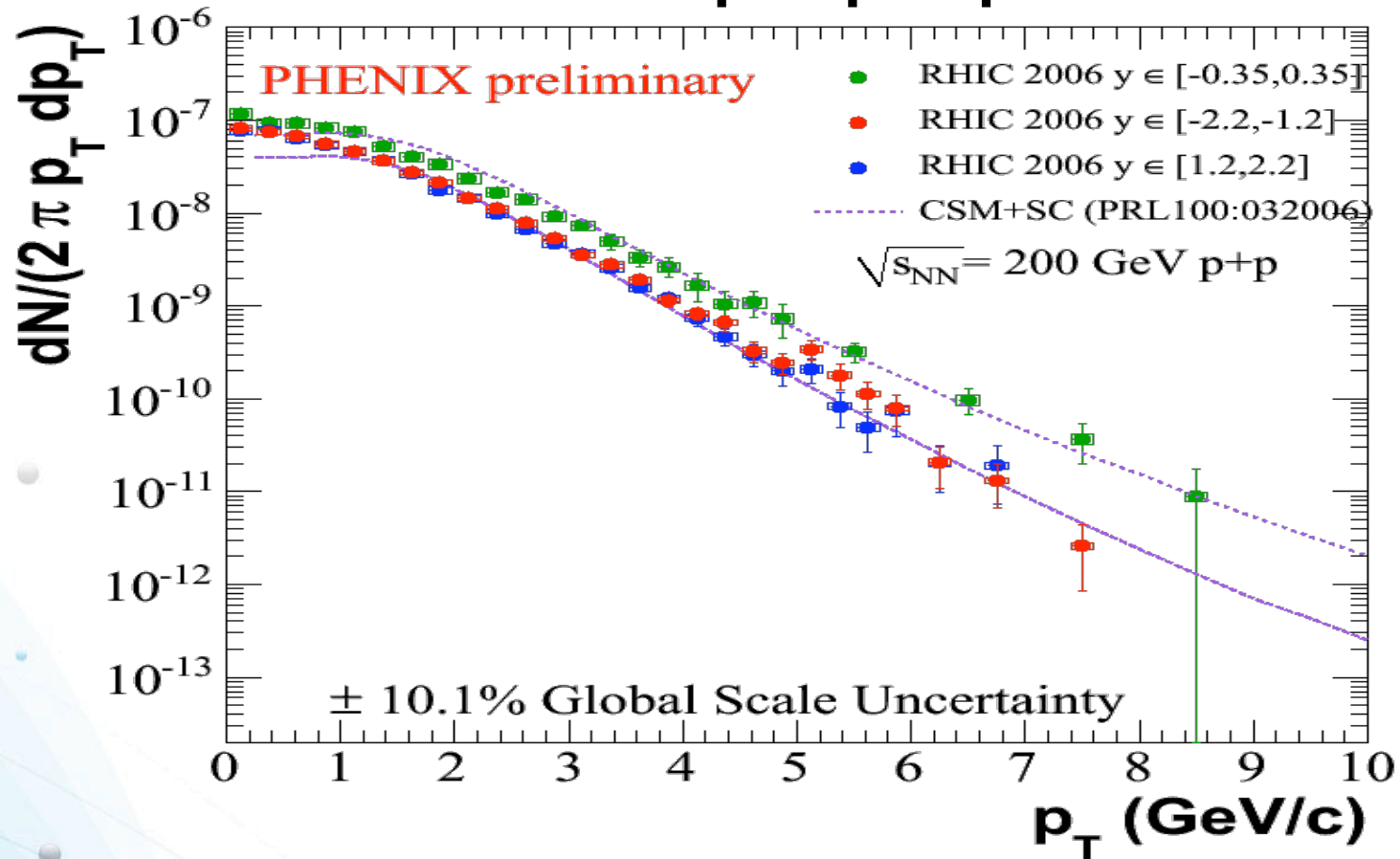
Talk by T Frawley

# RHIC 2006 p+p spectra



- ~3X increase in statistics ~2x in systematic over publication [PRL98:232002].
- Powerful data set to test production models.

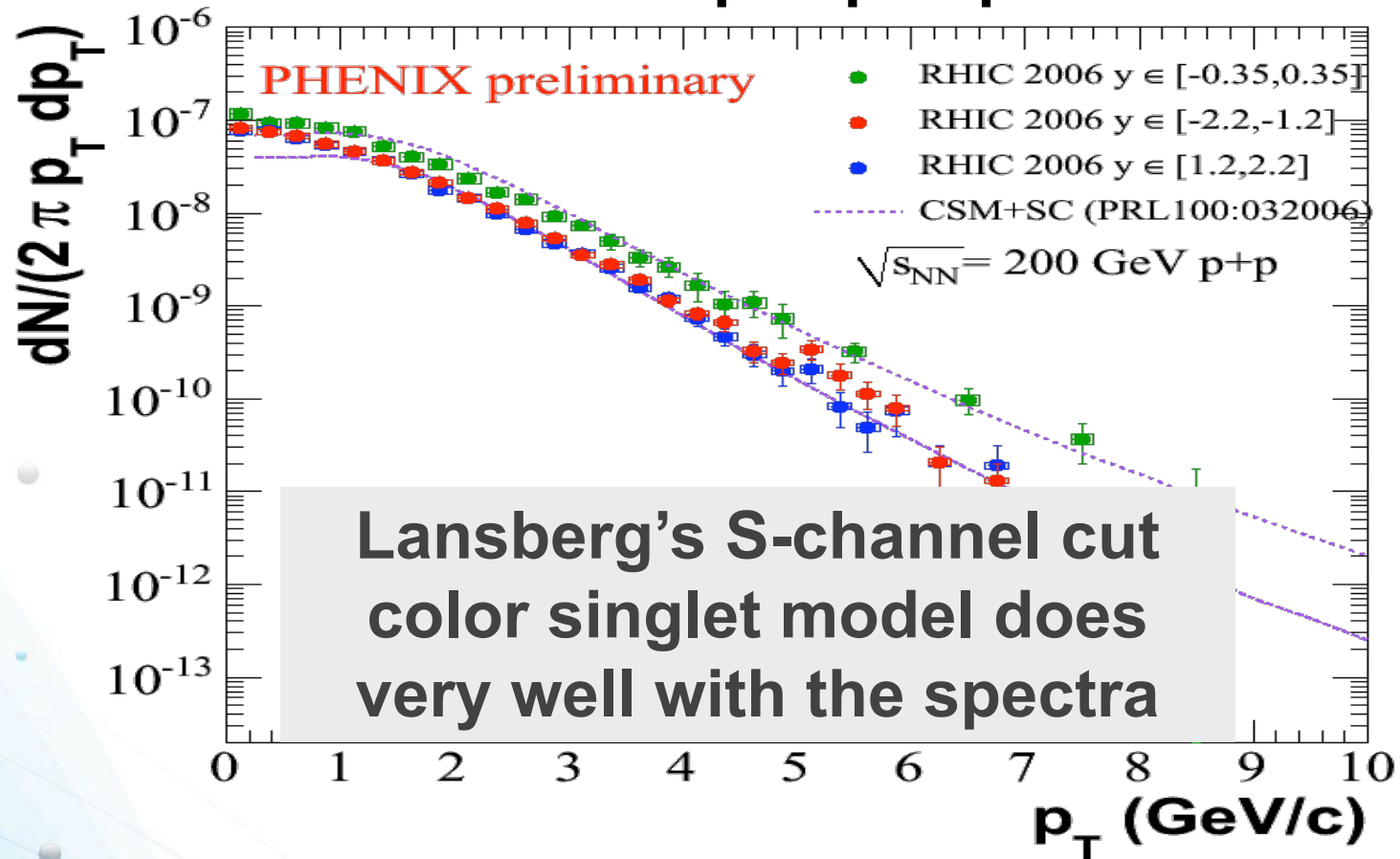
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# RHIC 2006 p+p spectra



- ~3X increase in statistics ~2x in systematic over publication [PRL98:232002].
- Powerful data set to test production models.

Talk by A. Linden Levy

# Hot Nuclear Matter Effects

**What are the in-medium quarkonium properties?**

**When does the  $J\psi$  melt?**

**Is there a  $J\psi$  peak in the spectral function?**

**What is the heavy quark potential?**

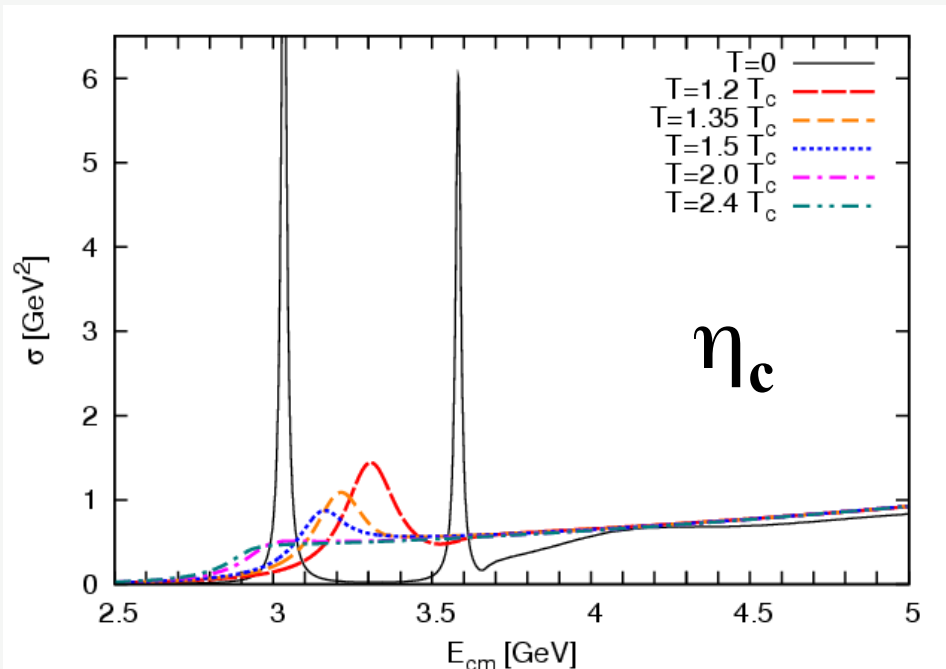
# Potential model approach for quarkonia at finite temperature

Spectral function calculations  
using some lattice-based potentials

**Talks by Ralf Rapp, Felix Riek, Agnes Mocsy**

## 2.4.2 Example from “Extended T-Matrix”

### S-Wave Spectral Function

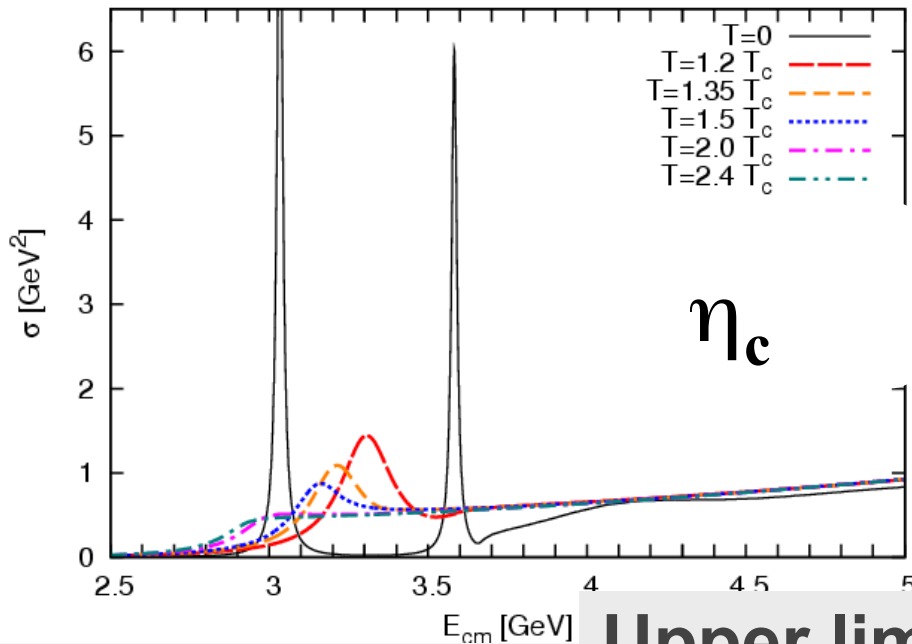


- **cc** propagator with  $\Gamma_c = 100$  MeV:
- S-wave “melting”  $T_{\text{diss}} \approx 1.5\text{-}2 T_c$
- correlator ratio temperature-stable

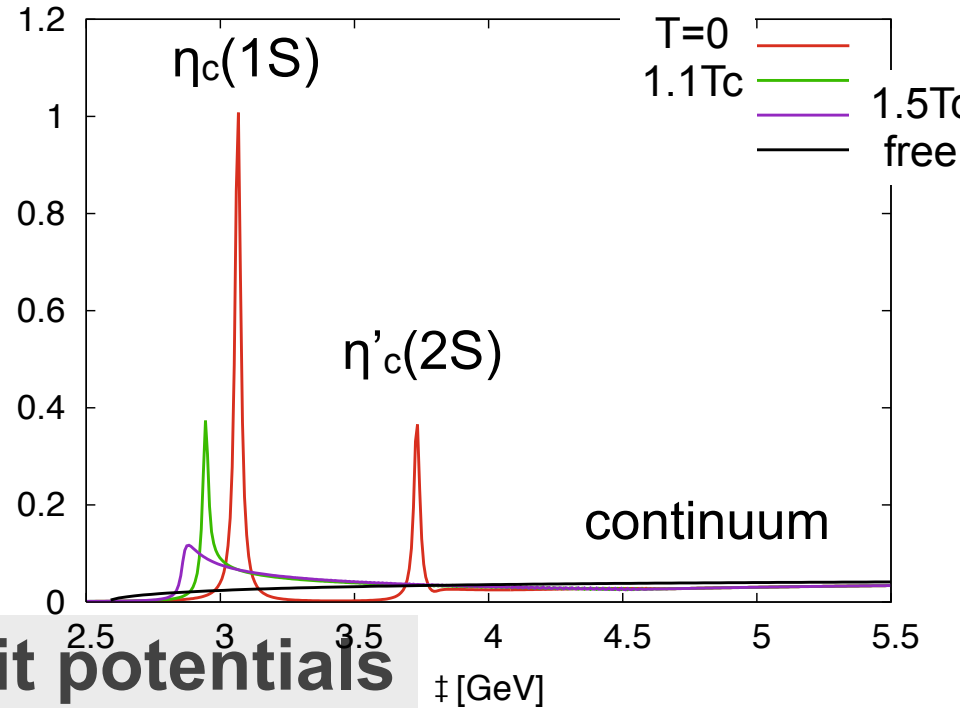
**Rapp, Riek**

# 2.4.2 Example from “Extended T-Matrix

## S-Wave Spectral Function



## Green's function calculation



## Upper limit potentials

- $cc$  propagator with  $\Gamma_c = 100 \text{ MeV}$ :
- S-wave “melting”  $T_{\text{diss}} \approx 1.5\text{-}2 T_c$
- correlator ratio temperature-stable

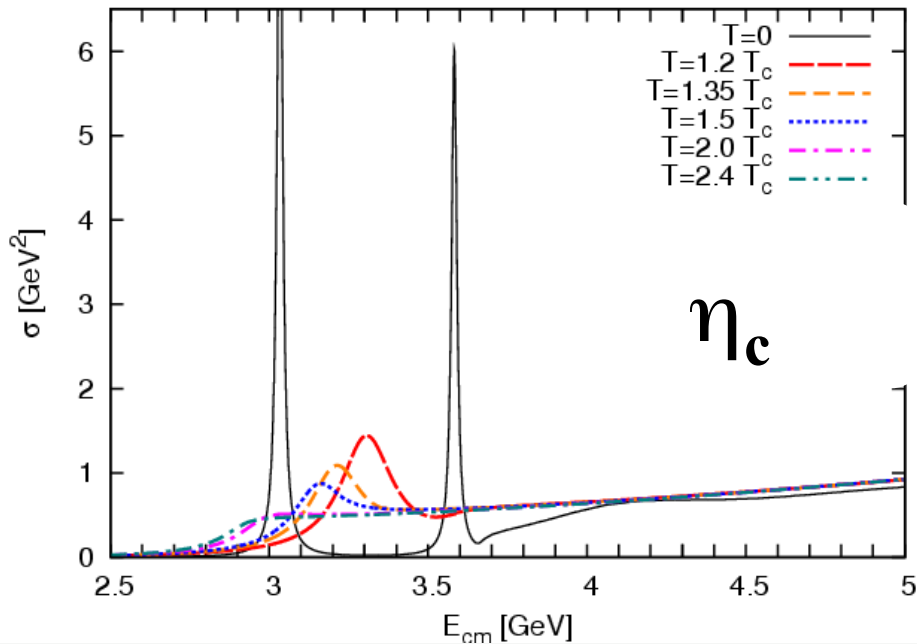
- $\Gamma_c = 13 \text{ MeV}$
- $T_{\text{diss}} \lesssim 1.2 T_c$
- correlator ratio agrees to 2%

Rapp, Riek

Mocsy, Petreczky

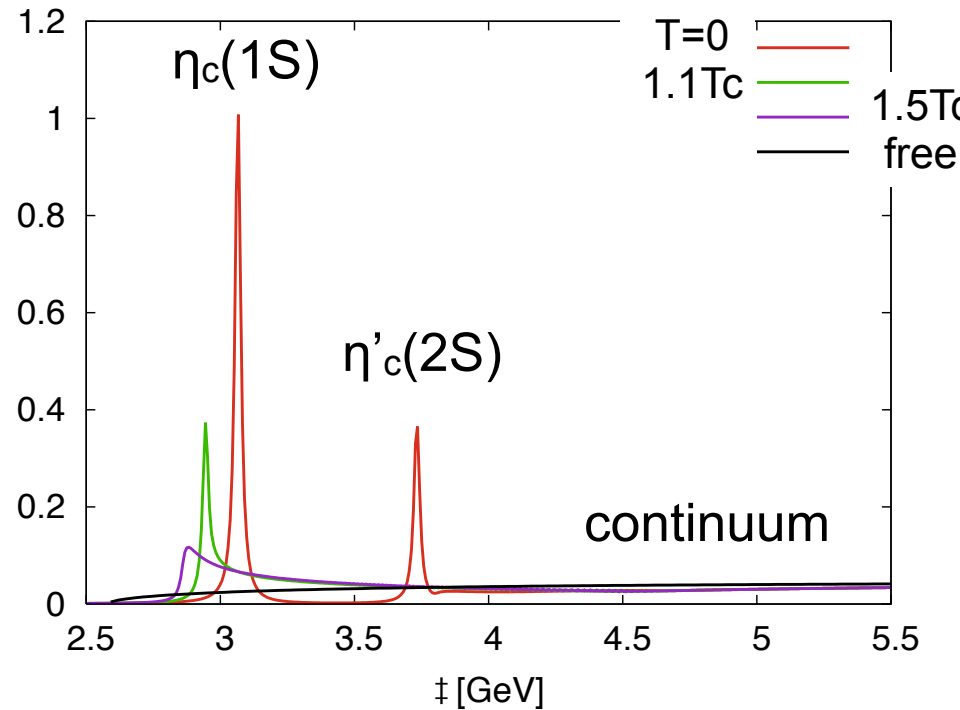
# Do We Agree or Disagree?

## S-Wave Spectral Function



- $cc$  propagator with  $\Gamma_c = 100$  MeV:
- S-wave “melting”  $T_{\text{diss}} \approx 1.5\text{--}2 T_c$
- correlator ratio temperature-stable

**Rapp, Riek**



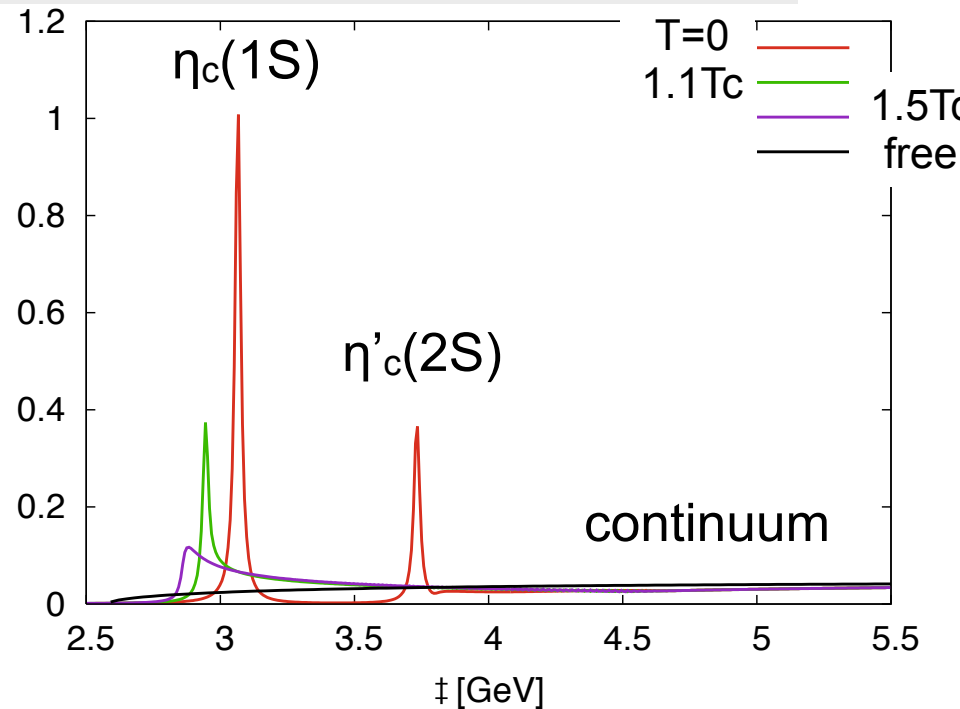
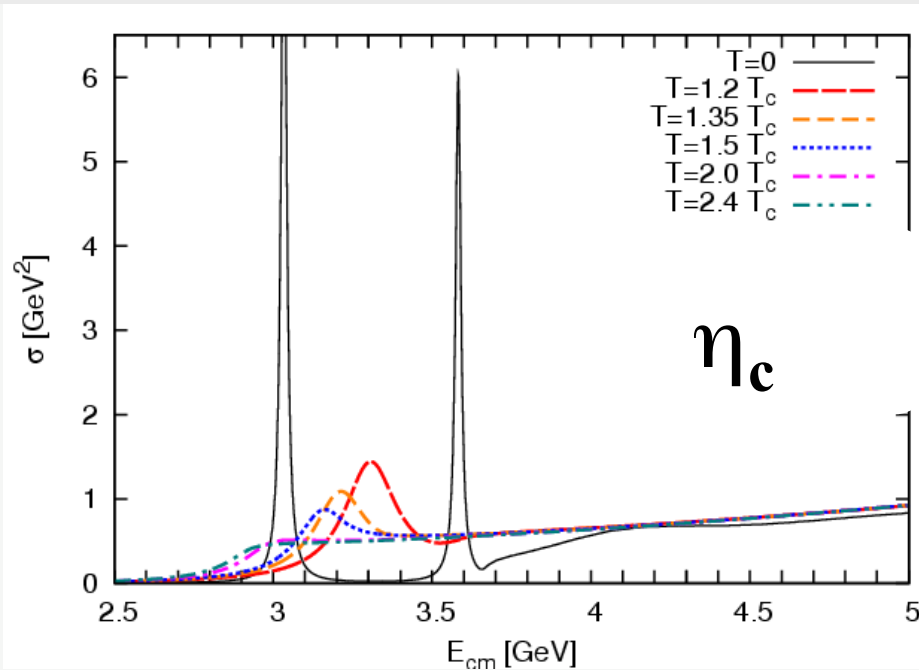
- $\Gamma_c = 13$  MeV
- $T_{\text{diss}} \lesssim 1.2 T_c$
- threshold enhancement
- correlator ratio agrees to 2%

**Mocsy, Petreczky**



# Do We Agree or Disagree?

My answer: **Yes, we agree.** Please have a careful look.



- cc propagator with  $\Gamma_c = 100$  MeV:
- S-wave “melting”  $T_{\text{diss}} \approx 1.5\text{-}2 T_c$
- correlator ratio temperature-stable

**Rapp, Riek**

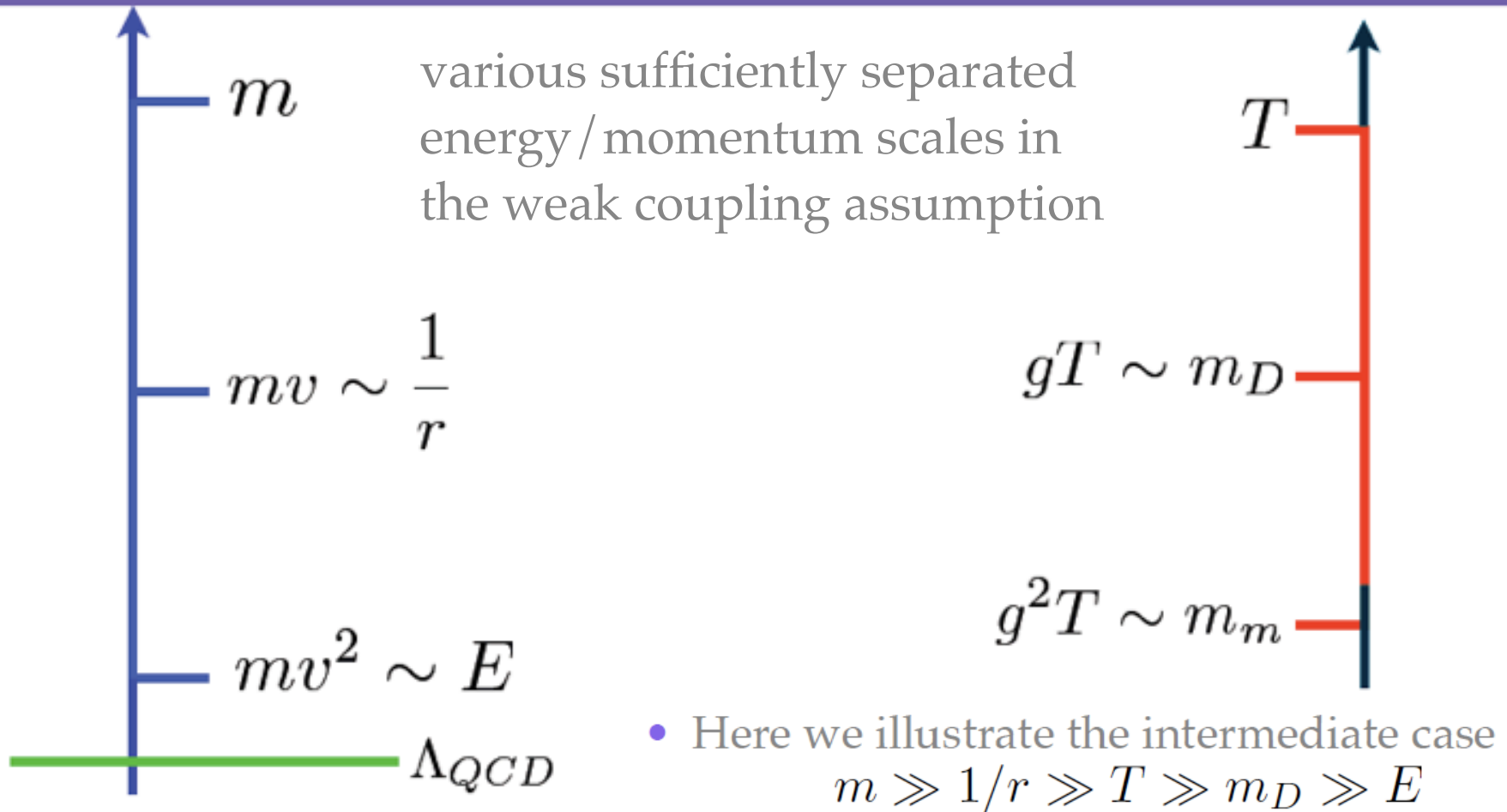
- $\Gamma_c = 13$  MeV
  - $T_{\text{diss}} \lesssim 1.2 T_c$
  - threshold enhancement
  - correlator ratio agrees to 2%
- Mocsy, Petreczky**

# The EFT approach for quarkonia

Systematic, non model-based  
QCD approach to the potential

**Talks by Jacopo Gighlieri, Andrea Beraudo**

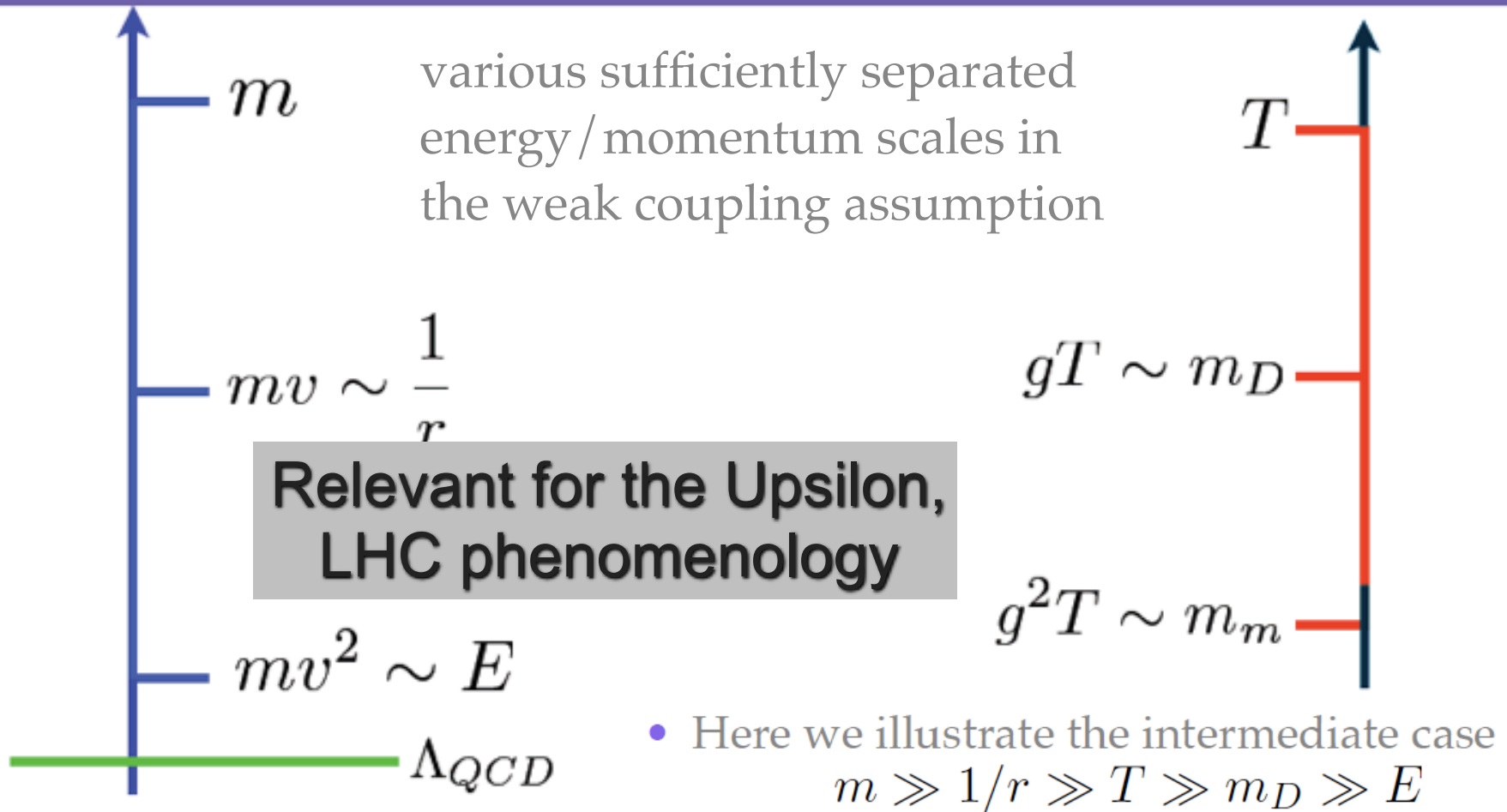
# Scales of the problem



Talk by J. Gighlieri

Brambilla JG Petreczky Vairo 2008

# Scales of the problem

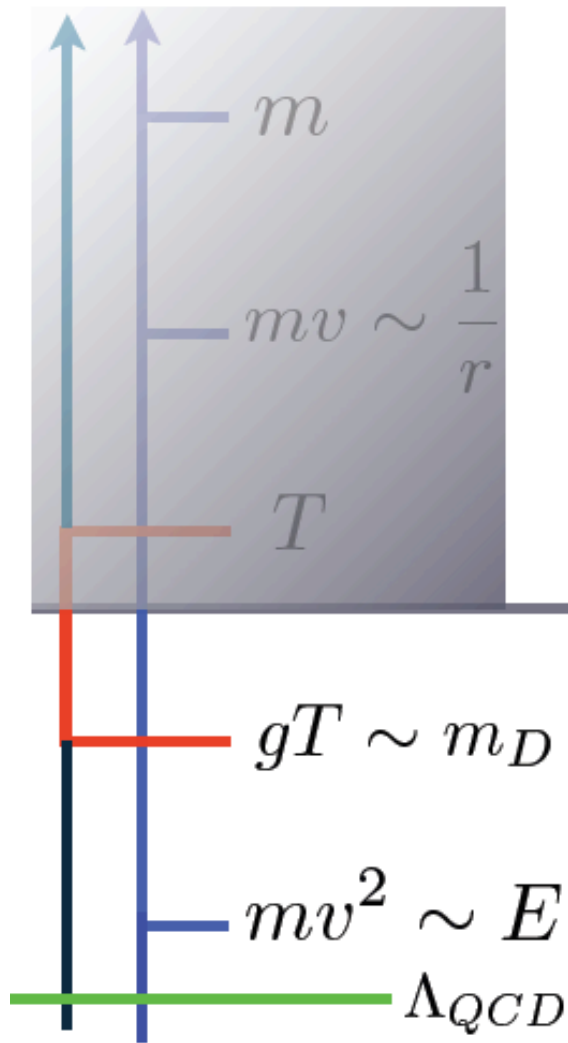


Talk by J. Gighlieri

Brambilla JG Petreczky Vairo 2008

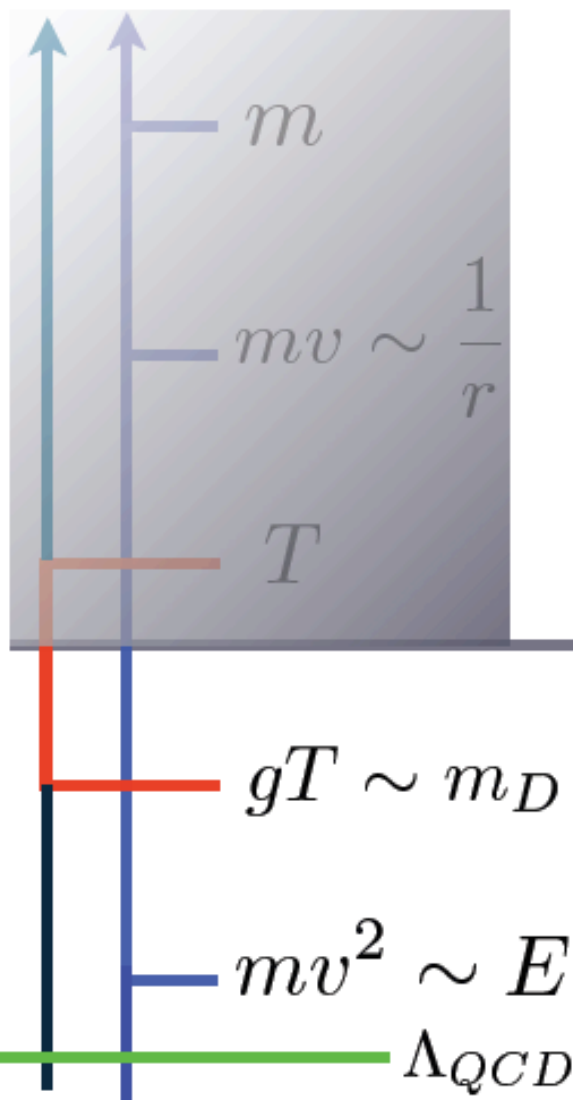
# The temperature

- First thermal corrections to the potential (power law)
- Corrections appear as loops in the effective theory
- Real and imaginary parts, contributing to energy and decay width observables

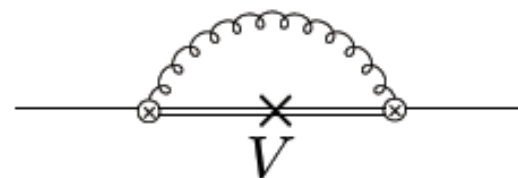


Talk by J. Gighlieri

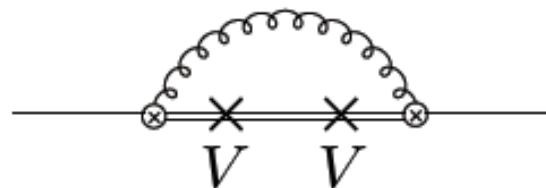
# Imaginary part can be as important as the real part



Talk by J. Gighlieri



$$\text{Re } \delta V_s(r) = \frac{\pi}{9} N_c C_F \alpha_s^2 r T^2 \quad \sim g^2 r^2 T^3 \times \frac{V}{T}$$



$$\text{Im } \delta V_s(r) = -\frac{N_c^2 C_F}{6} \alpha_s^3 T \quad \sim g^2 r^2 T^3 \times \left(\frac{V}{T}\right)^2$$

- The imaginary part correspond to singlet-to-octet thermal breakup

# In-medium heavy-quark spectral function: a path-integral approach

**Andrea Beraudo**

*University of Torino and CERN-Th.Div. (Centro-Fermi fellowship)*

BNL, 14<sup>th</sup> – 18<sup>th</sup> December 2009

*Work in progress in collaboration with J.P. Blaizot (CEA-Saclay),  
G. Garberoglio and P. Faccioli (University of Trento)*

Nucl. Phys. A 830, 319C-322C (2009)

**Talk by A. Beraudo**



This method will be really handy when we have to solve the EFT

## In-medium heavy-quark spectral function: a path-integral approach

Andrea Beraudo

### Our goal

We wish to perform a study resulting

- numerically less expensive than lattice calculations (hence allowing a more robust reconstruction of the spectral function);
- capable to get a deeper physical insight on the processes involved.

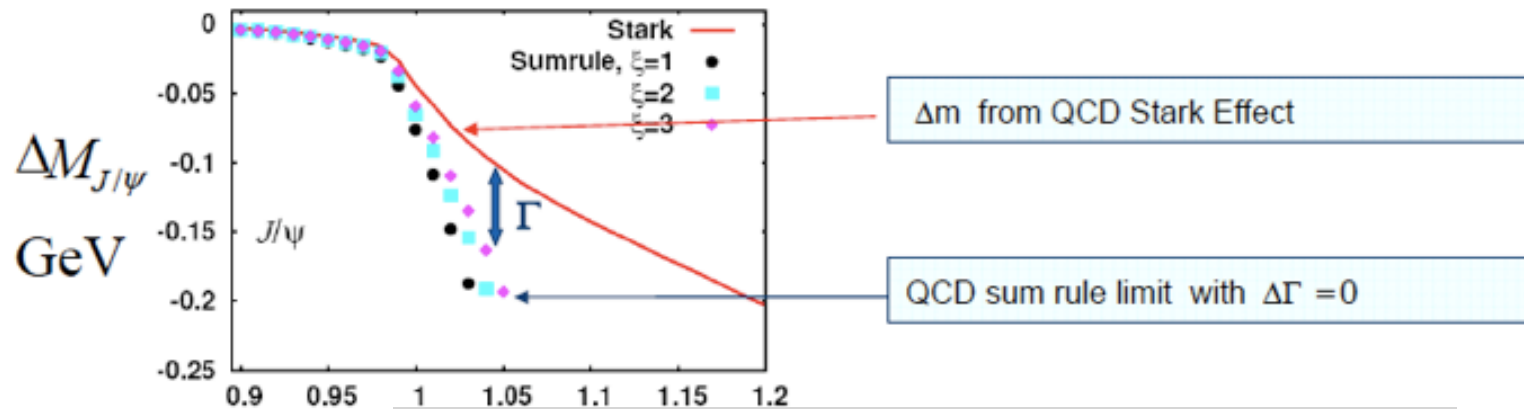
**Talk by A. Beraudo**

# QCD sum-rule approach for quarkonia at finite temperature

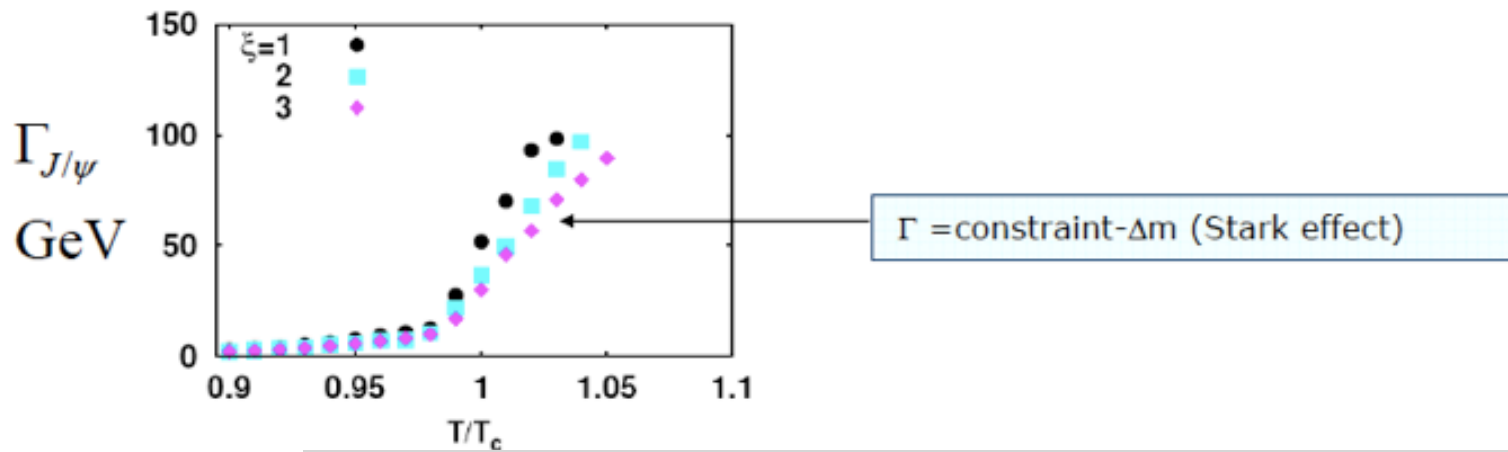
Study of heavy quark system very near  $T_c$

**Talk by Su Houng Lee**

## Mass and width of $J/\psi$ near $T_c$ (Morita, Lee 08)

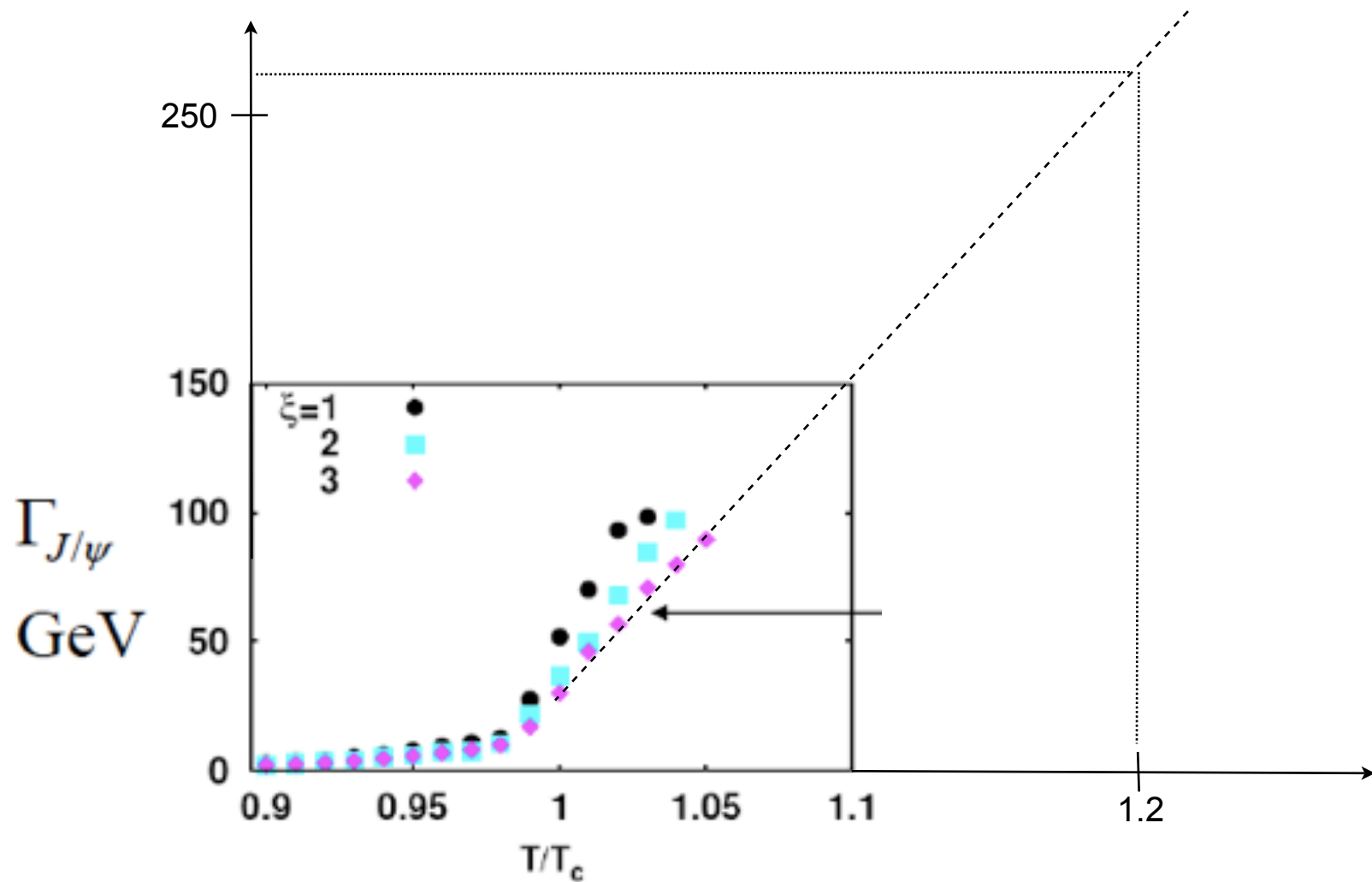


Onset of abrupt mass shift at  $T_c$



Sudden increase in the width near  $T_c$

Please allow me, just for fun ...



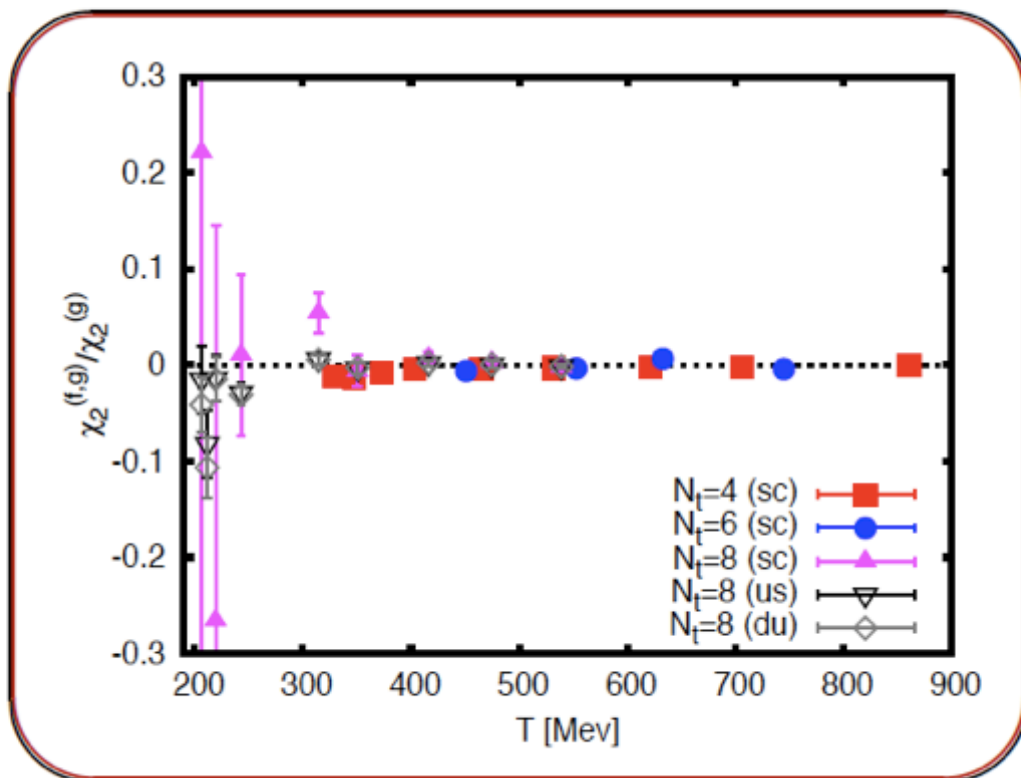
# Lattice QCD results for heavy quarks at finite temperature

Heavy quark susceptibilities, correlations

**Talk by Swagato Mukherjee**

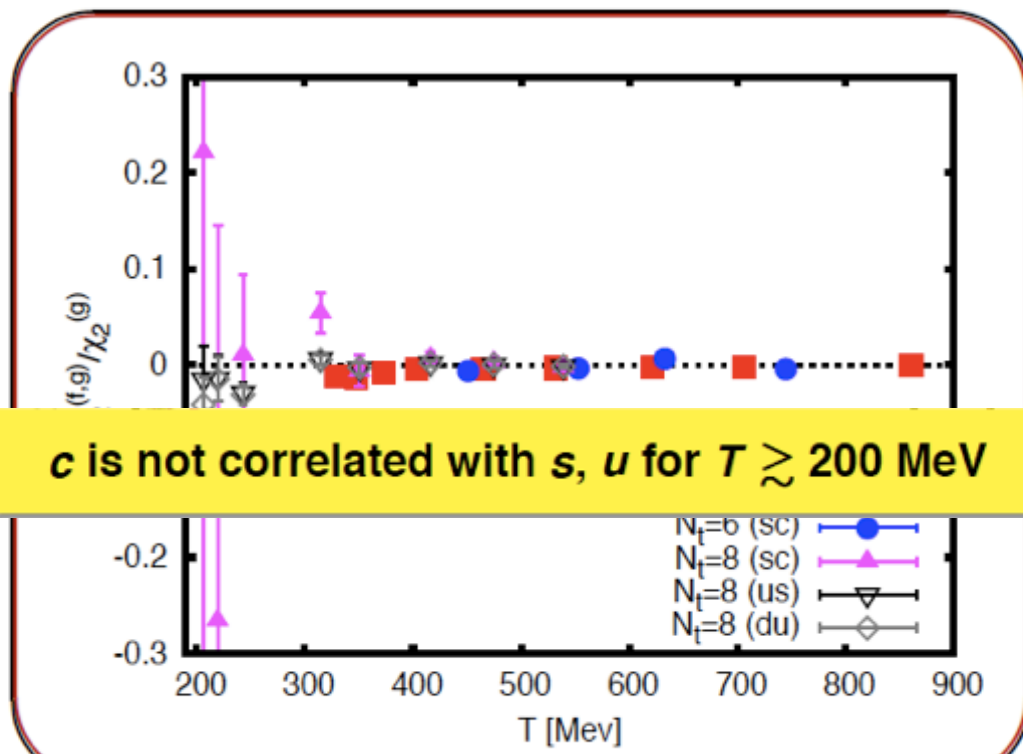
## Flavor correlations

$$\chi_2^{(f,g)} / \chi_2^{(g)} = \langle \mathcal{N}_f \mathcal{N}_g \rangle / \langle \mathcal{N}_g \mathcal{N}_g \rangle$$



## Flavor correlations

$$\chi_2^{(f,g)} / \chi_2^{(g)} = \langle \mathcal{N}_f \mathcal{N}_g \rangle / \langle \mathcal{N}_g \mathcal{N}_g \rangle$$



Implies that heavy-light mesons do not exist above  $T_c$

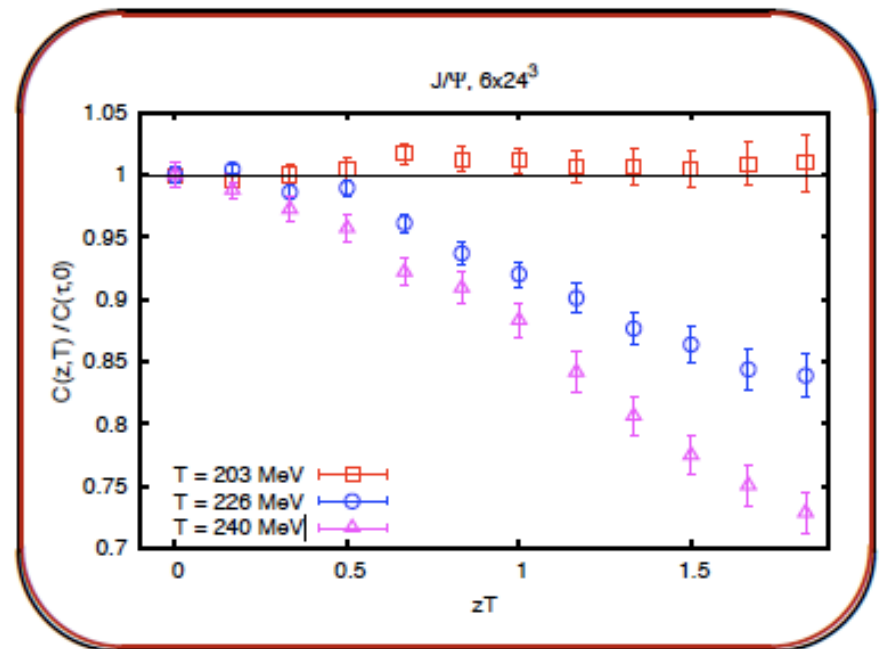
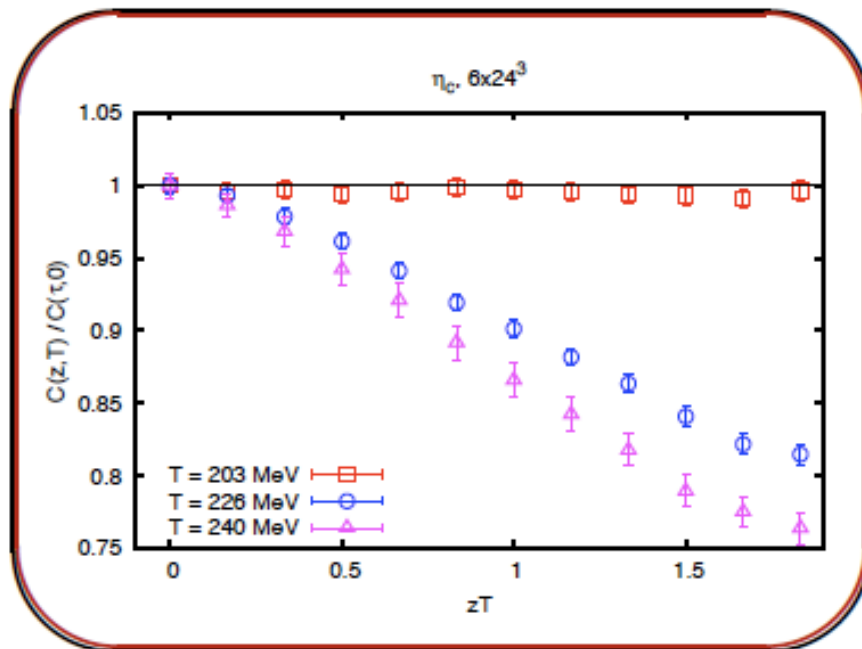
*Also: Charm number fluctuations compatible with free quark gas with slightly  $T$ -dependent mass*



# Spatial correlations of charmonia

$$T \rightarrow 0 \therefore M = m_{had}$$

(Spatial correlator at T)/(Spatial correlator T=0)



Possible indication that there are no quarkonium states

# Bridging to experimental data

## Through a dynamical model

**Talks by Clint Young, Su Houng Lee, Ralf Rapp**

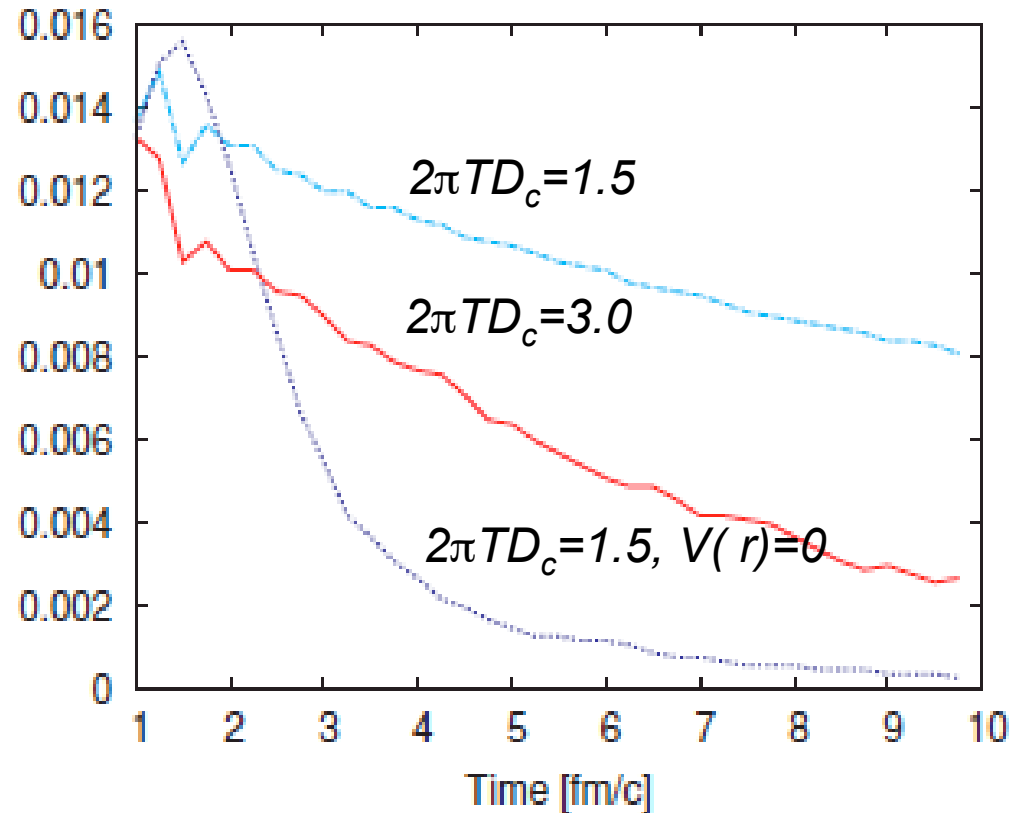
**If we take seriously that the Jpsi melts,  
broadens, not there, ...  
why don't we see 100% suppression?**

# General features of the Langevin-with-interaction model for charmonium

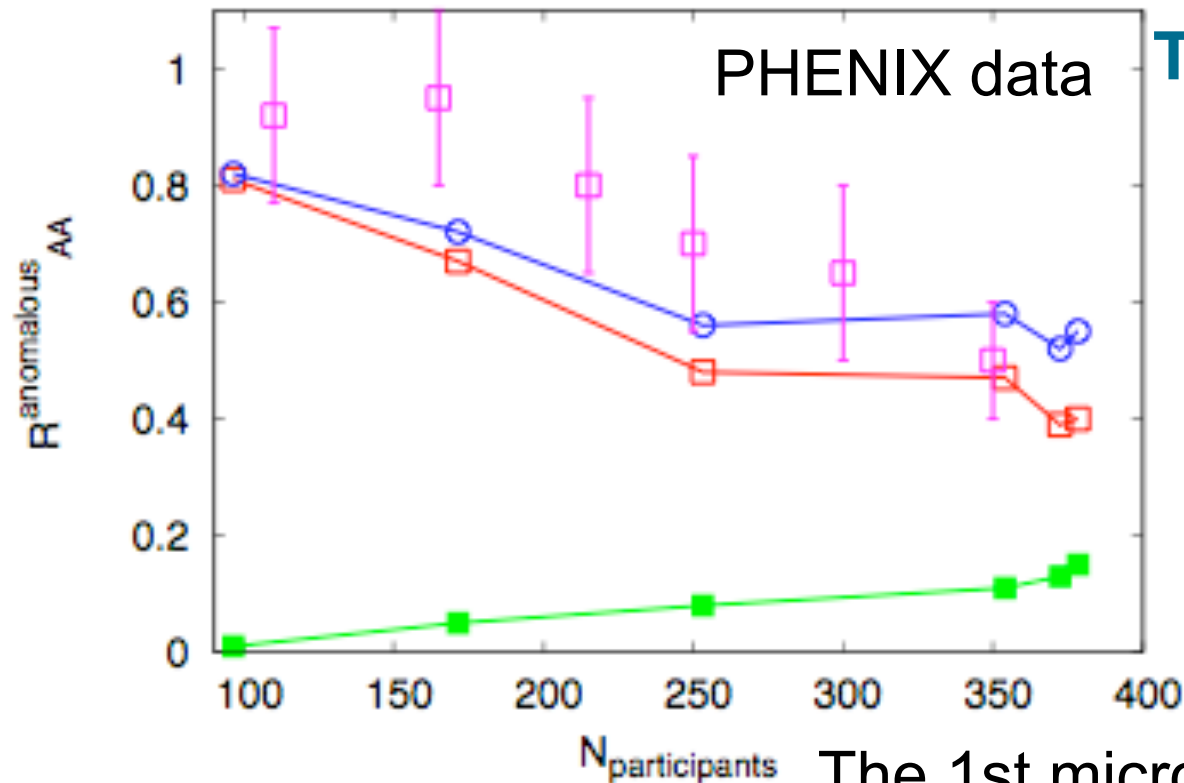
$\frac{\text{\# of correlated pairs}}{\text{total \# of c-cbar pairs}}$

Talk by C Young

- Langevin + hydro (the c-cbar pairs move through the sQGP - drag, Langevin, potential)
- Medium is trying to prevent quarks from being separated
- An ensemble evolving with this large drag coefficient equilibrates rapidly in momentum space, and then experiences slow spatial diffusion.
- Timescales become very important to determine the change in yields of charmonium due to a deconfined phase.



# $R_{AA}$ for $T_c=190$ MeV



Talk by C Young

correlated pairs

The 1st microscopic calculation of non-correlated recombination!

*The Langevin-with-interaction model for charmonium in sQGP can explain the suppression pattern for the  $J/\psi$  in Au+Au collisions*

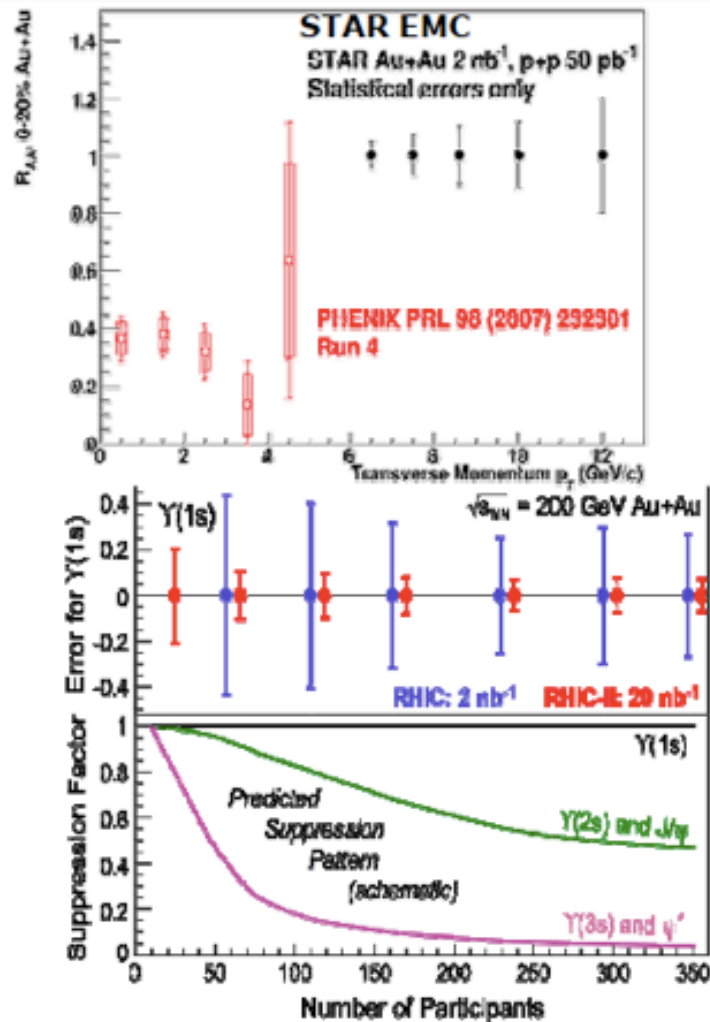
# The future is bright

## Experimental upgrades in the pipeline

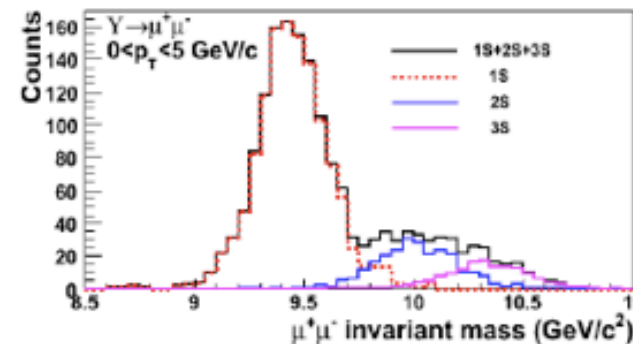
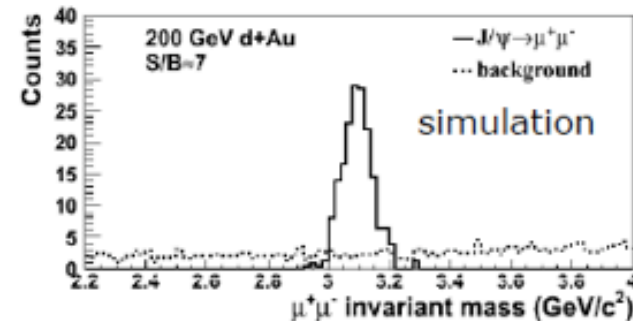
**Talk by Zhangbu Xu**

# The future is bright

## High luminosity for $\Upsilon$ & $J/\psi$



L.Ruan et al., 0904.3774, JPG36(2009);  
Z. Xu, BNL LDRD project 07-007  
**STAR Muon Telescope Detector**



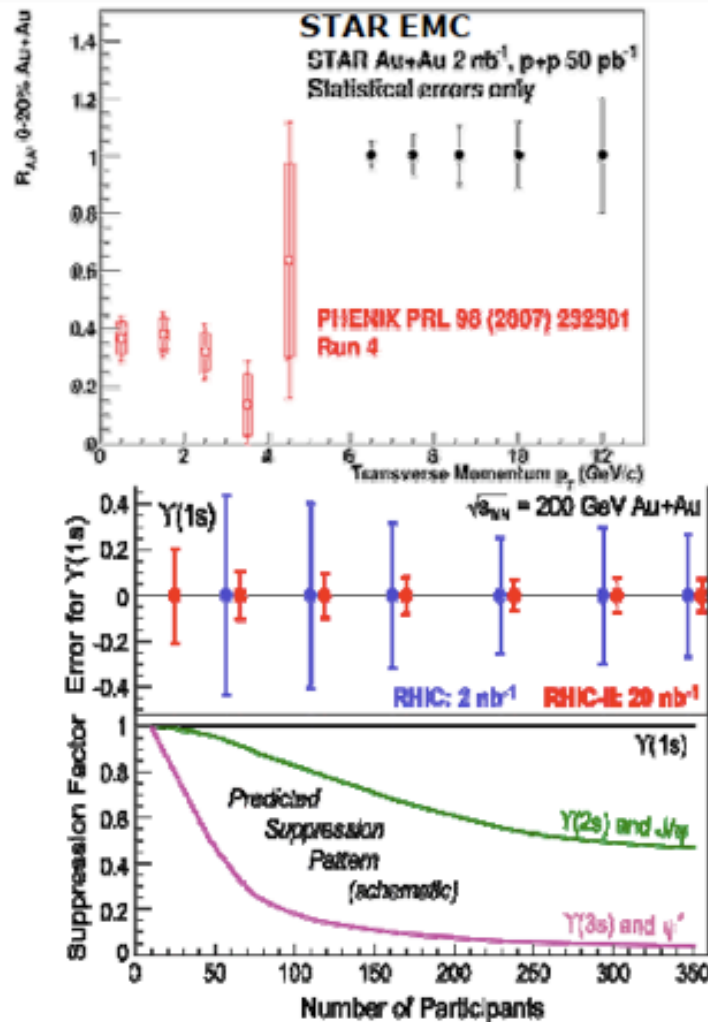
**MTD: excellent mass resolution for  $\tau$   
separate different  $\tau$  states**

14

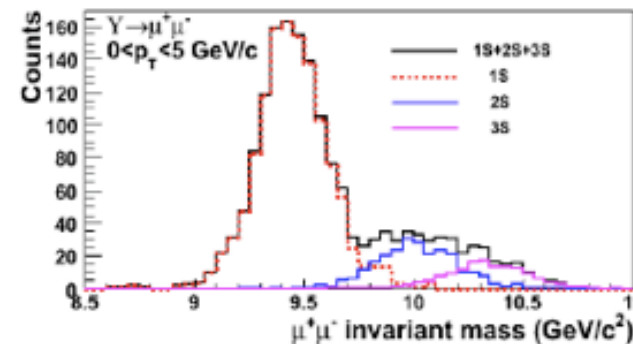
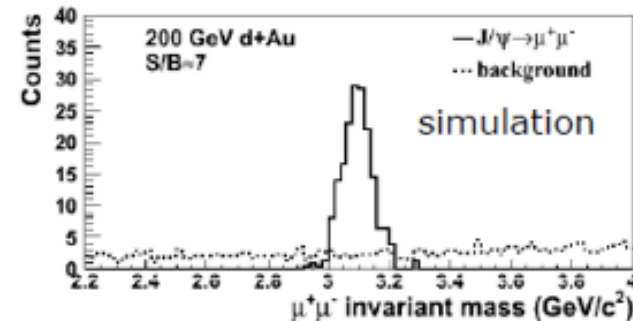
**Talk by Z. Xu**

# The future is bright

## High luminosity for $\Upsilon$ & $J/\psi$



L.Ruan et al., 0904.3774, JPG36(2009);  
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**MTD: excellent mass resolution for  $\Upsilon$   
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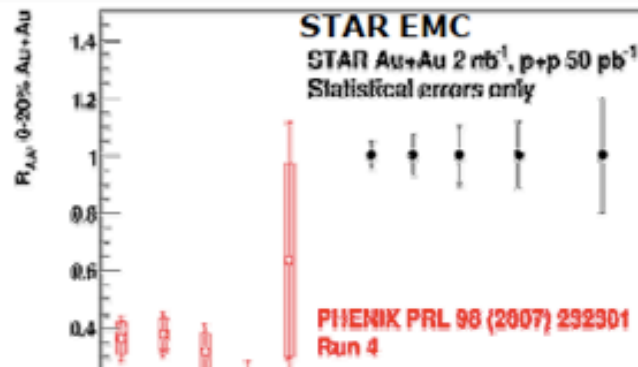
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**Talk by Z. Xu**

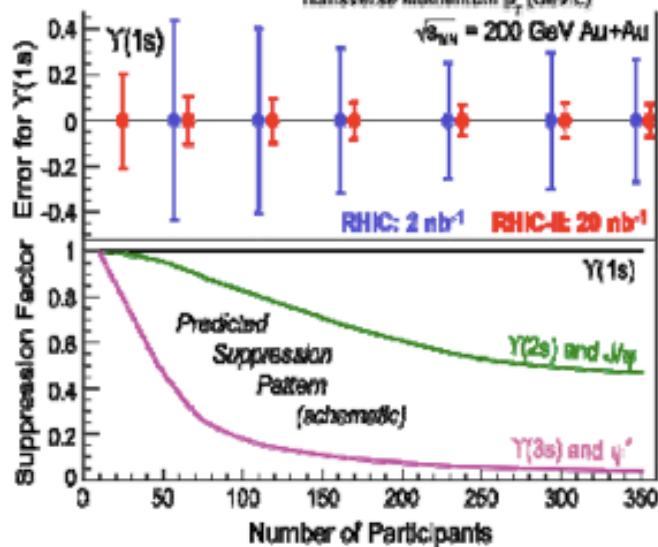


# The future is bright

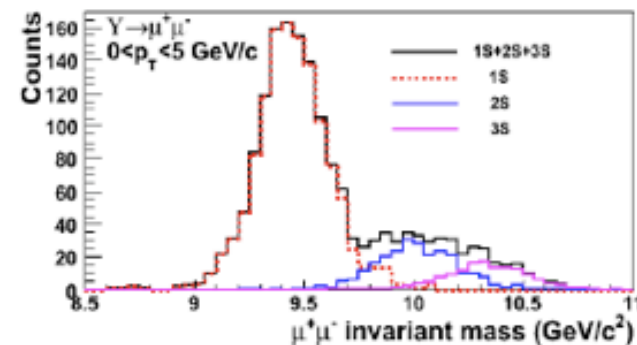
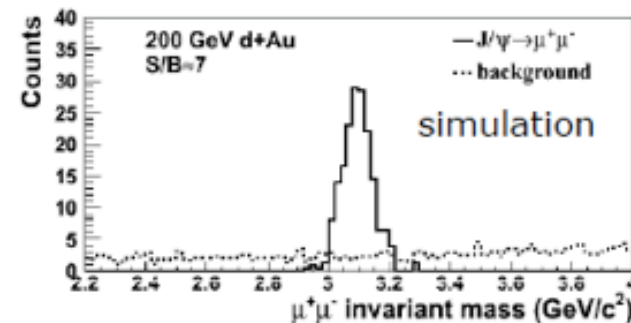
## High luminosity for $\Upsilon$ & $J/\psi$



Factor 10 in luminosity



L.Ruan et al., 0904.3774, JPG36(2009);  
Z. Xu, BNL LDRD project 07-007  
STAR Muon Telescope Detector



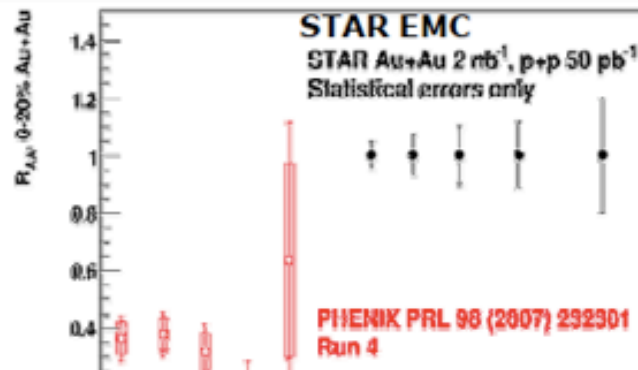
MTD: excellent mass resolution for  $\Upsilon$   
separate different  $\Upsilon$  states

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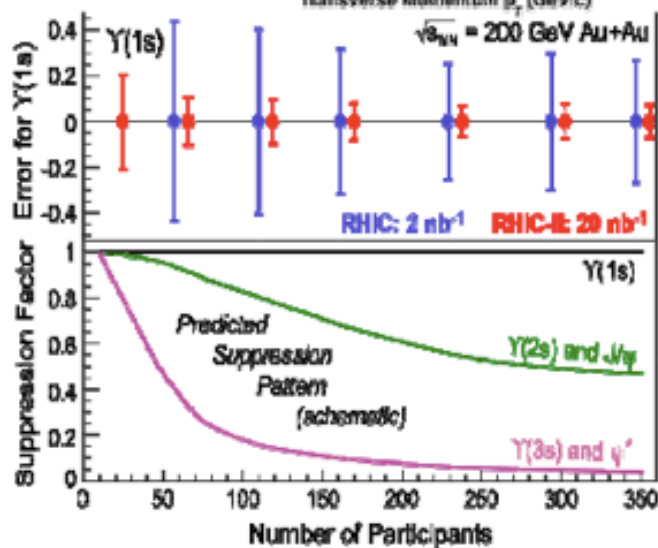
Talk by Z. Xu

# The future is bright

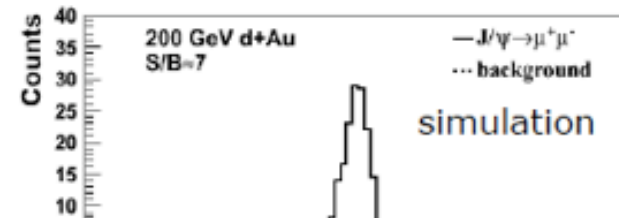
## High luminosity for $\Upsilon$ & $J/\psi$



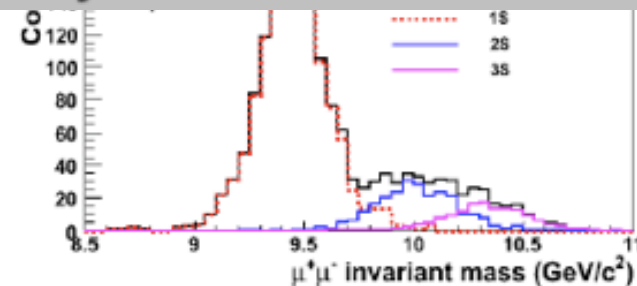
Factor 10 in luminosity



L.Ruan et al., 0904.3774, JPG36(2009);  
Z. Xu, BNL LDRD project 07-007  
STAR Muon Telescope Detector



Clear ground state peaks  
Maybe even excited ones?!



MTD: excellent mass resolution for  $\Upsilon$   
separate different  $\Upsilon$  states

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Talk by Z. Xu

**THE END**